

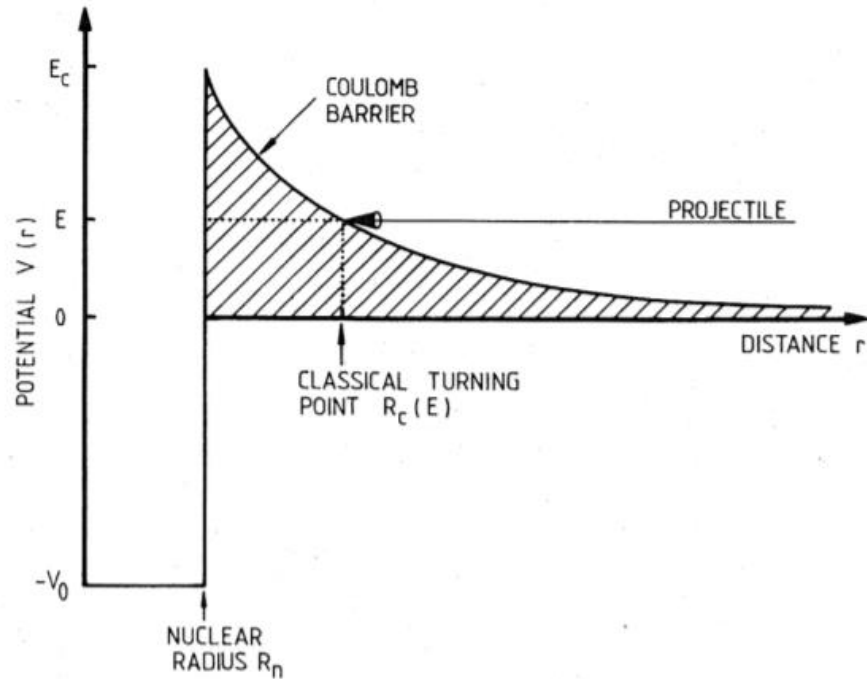
A Nap energiatermelése

Magfizika szeminárium 2019. 10. 03.

Marx Pál Fülöp

Coulomb gát

$$V_C(r) = \frac{Z_1 Z_2 e^2}{r},$$



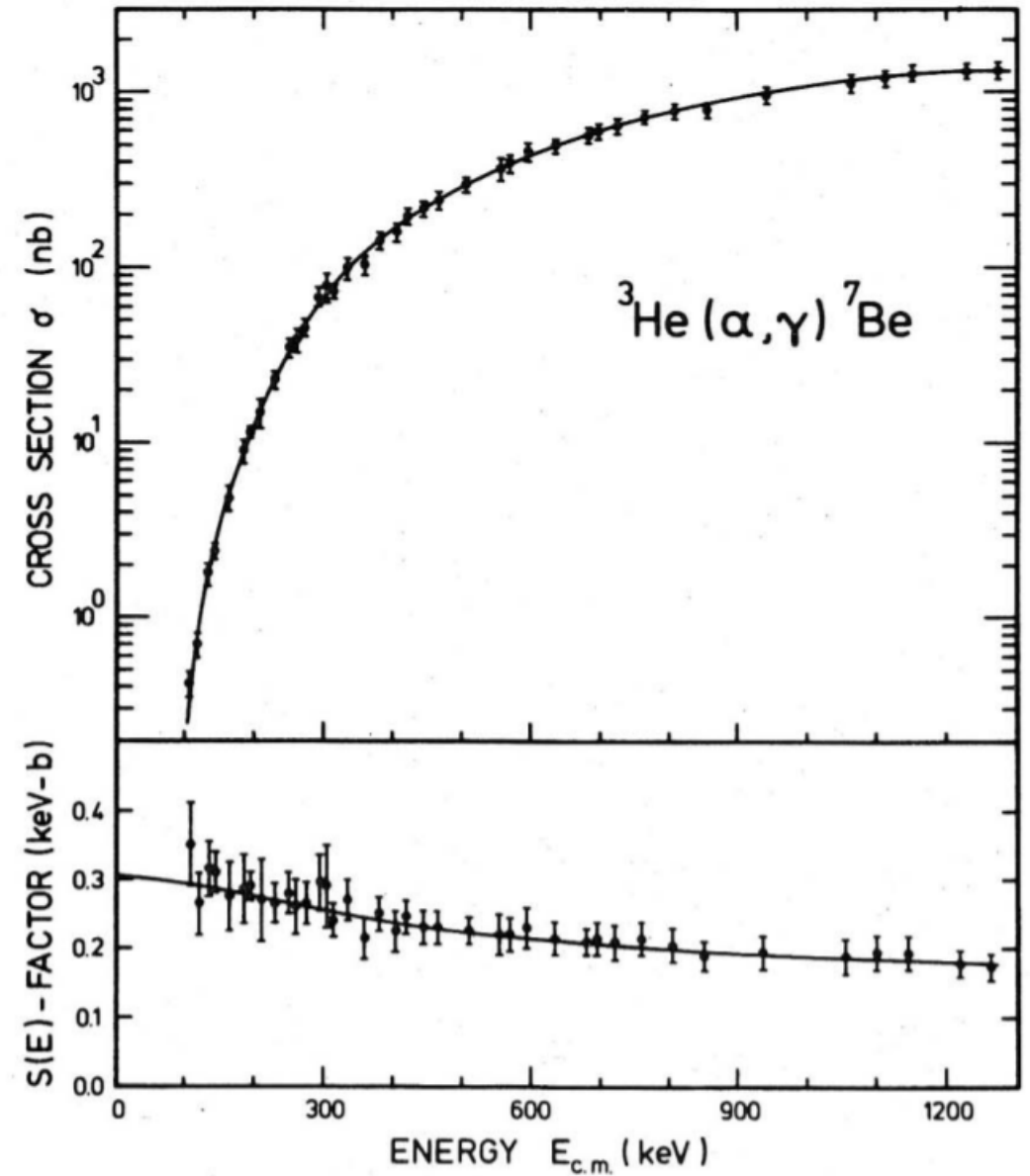
$$P = \exp \left\{ -2 \left[\frac{2\mu}{\hbar^2} (E_C - E) \right]^{\frac{1}{2}} R_C \left[\frac{\arctan \left(\frac{R_C}{R_n} - 1 \right)^{\frac{1}{2}}}{\left(\frac{R_C}{R_n} - 1 \right)^{\frac{1}{2}}} - \frac{R_n}{R_C} \right] \right\}$$

TABLE 4.1 Tunneling Probability for the Reaction $p + p$ as a Function of Center-of-Mass Energy E

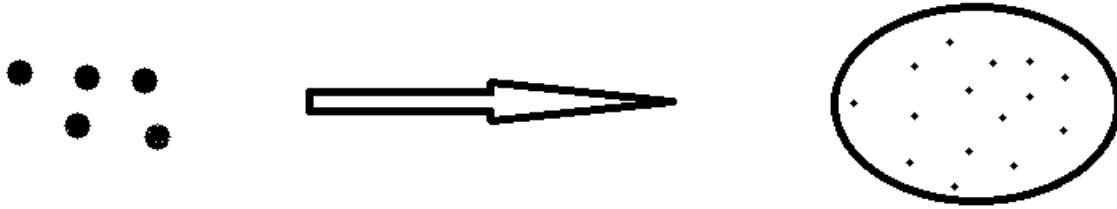
E (keV)	Tunneling Probability	
	Exact Expression	Approximation
1	8.9×10^{-10}	2.5×10^{-10}
2	5.6×10^{-7}	1.6×10^{-7}
5	1.7×10^{-4}	5.0×10^{-5}
10	3.1×10^{-3}	9.2×10^{-4}
20	2.4×10^{-2}	7.1×10^{-3}
50	0.14	4.4×10^{-2}
100	0.35	0.11
200	0.64	0.21
500	0.99	0.37
550 ^a	1.00	0.39

Coulomb gát

$$\langle \sigma v \rangle = \left(\frac{8}{\pi \mu} \right)^{1/2} \frac{1}{(kT)^{3/2}} \int_0^\infty S(E) \exp \left[-\frac{E}{kT} - \frac{b}{E^{1/2}} \right] dE$$



Reakcióráta



- Lövedék áramsűrűsége: $\frac{N}{A dt} = \frac{N ds}{A ds dt} = n_x v$
- Reakcióráta: $n_x n_y \langle \sigma(v) v \rangle$

Hőmérsékleti eloszlás

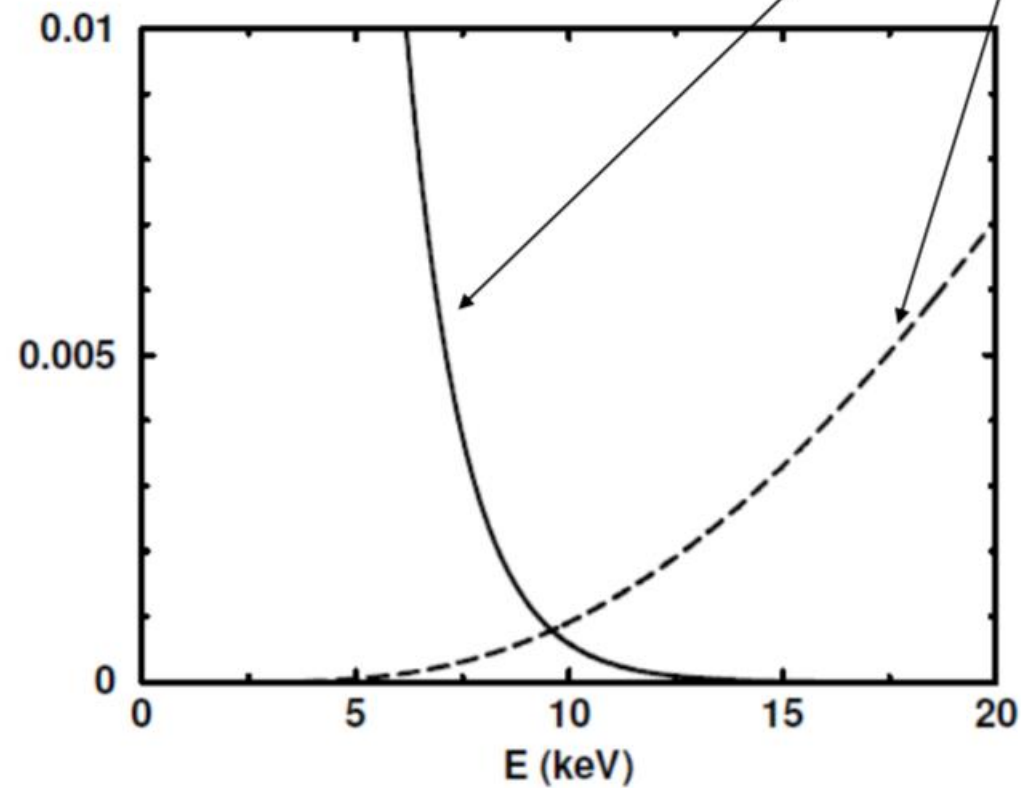
$$\left(\frac{\mu}{2\pi kT}\right)^{3/2} 4\pi \exp\left(-\frac{\mu v^2}{2kT}\right) v^2 dv \quad \frac{1}{(kT)^{3/2}} \frac{2}{\sqrt{\pi}} \exp\left(-\frac{E}{kT}\right) \sqrt{E} dE$$

Termikus reakcióráta

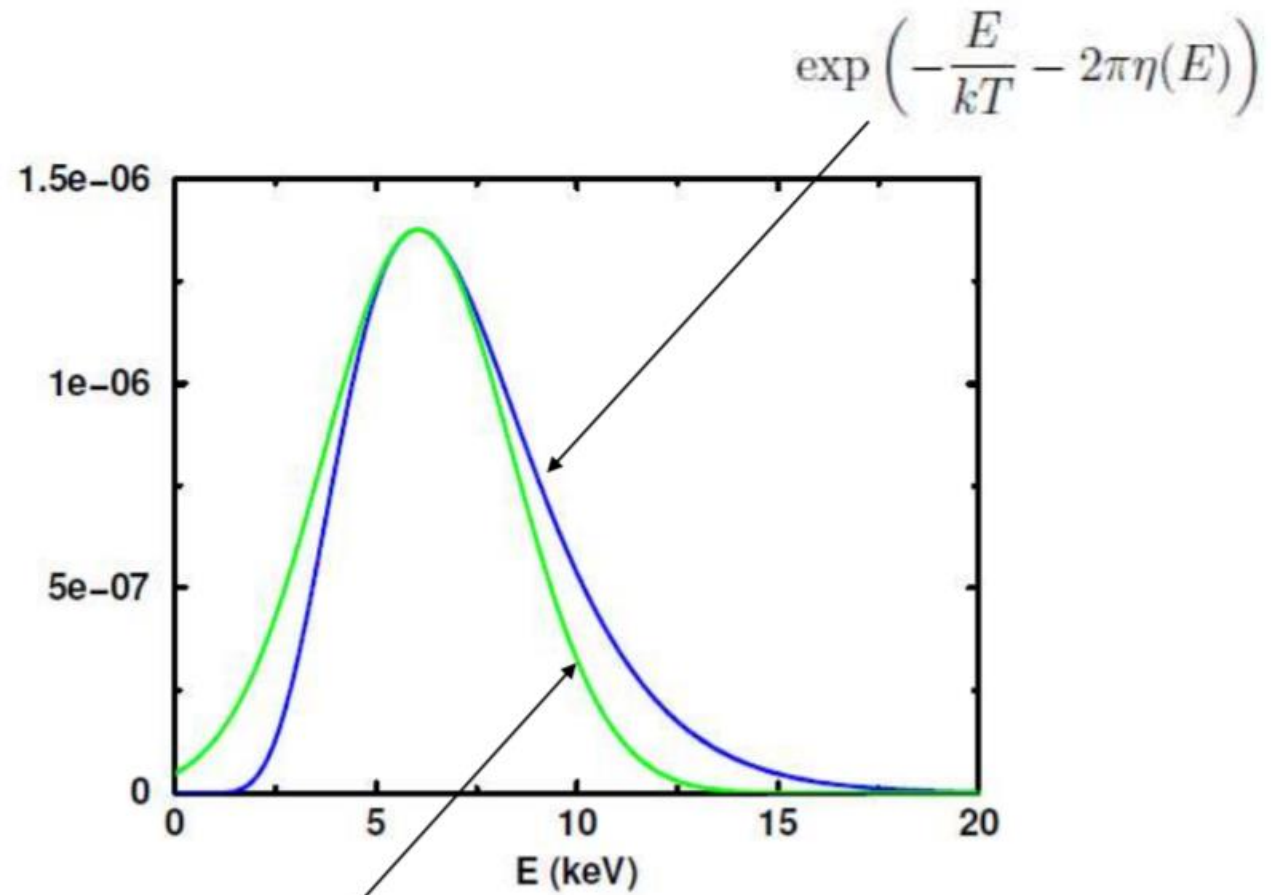
$$\begin{aligned} n_A n_B \langle \sigma v \rangle &= n_A n_B \left(\frac{8}{\pi\mu}\right)^{1/2} \frac{1}{(kT)^{3/2}} \int_0^\infty E \exp\left(-\frac{E}{kT}\right) \sigma(E) dE \\ &= n_A n_B \left(\frac{8}{\pi\mu}\right)^{1/2} \frac{1}{(kT)^{3/2}} \int_0^\infty \exp\left(-\frac{E}{kT} - 2\pi\eta(E)\right) S(E) dE \end{aligned}$$

Termikus reakcióráta

$$\begin{aligned}n_A n_B \langle \sigma v \rangle &= n_A n_B \left(\frac{8}{\pi \mu} \right)^{1/2} \frac{1}{(kT)^{3/2}} \int_0^\infty E \exp\left(-\frac{E}{kT}\right) \sigma(E) dE \\ &= n_A n_B \left(\frac{8}{\pi \mu} \right)^{1/2} \frac{1}{(kT)^{3/2}} \int_0^\infty \exp\left(-\frac{E}{kT} - 2\pi\eta(E)\right) S(E) dE\end{aligned}$$



Gamow-csúcs



$$I_{\max} \exp\left(-\frac{(E - E_0)^2}{(\Delta/2)}\right) \text{ Gauss közelítés}$$

Coulomb gát

$$p + p: \quad E_0 = 5.9 \text{ keV ,}$$

$$p + {}^{14}\text{N}: \quad E_0 = 26.5 \text{ keV ,}$$

$$\alpha + {}^{12}\text{C}: \quad E_0 = 56 \text{ keV ,}$$

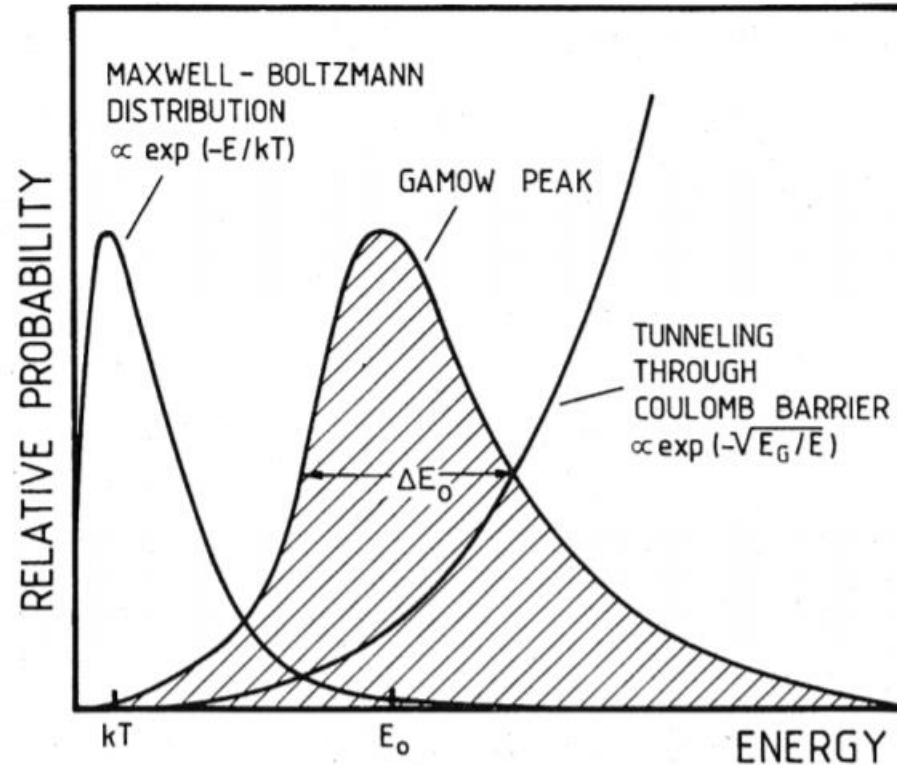
$${}^{16}\text{O} + {}^{16}\text{O}: \quad E_0 = 237 \text{ keV .}$$

$$p + p: \quad I_{\text{max}} = 1.1 \times 10^{-6} ,$$

$$p + {}^{14}\text{N}: \quad I_{\text{max}} = 1.8 \times 10^{-27} ,$$

$$\alpha + {}^{12}\text{C}: \quad I_{\text{max}} = 3.0 \times 10^{-57} ,$$

$${}^{16}\text{O} + {}^{16}\text{O}: \quad I_{\text{max}} = 6.2 \times 10^{-239} .$$



Csillagok szerkezete

A csillag belső szerkezete függ a csillag tömegétől!

$$E_* = L * T \approx M$$

$$L \approx M^{3.5}$$

$$T \approx M^{-2.5}$$

Heat Transfer of Stars

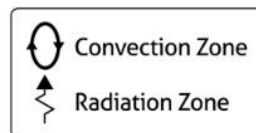
> 1.5 solar masses



0.5 - 1.5 solar masses



< 0.5 solar masses



A csillagok egyenletei:

$$\frac{dP}{dr} = -\frac{Gm\rho}{r^2},$$

$$\frac{dm}{dr} = 4\pi r^2 \rho.$$

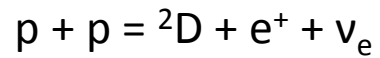
$$\frac{dl}{dr} = 4\pi r^2 \rho(\epsilon - \epsilon_\nu),$$

$$\frac{dT}{dr} = -\frac{1}{k} \frac{l}{4\pi r^2},$$

$$\frac{dT}{dr} = -\frac{3\kappa\rho l}{64\pi r^2 \sigma T^3},$$

- Sűrűség ρ
- Hőmérséklet T
- Teljes nyomás (anyag és sugárzás) P
- Luminozitás l
- Energiatermelési ráta ϵ
- Hővezető képesség k
- Opacitás κ
- Stefan-Boltzmann állandó σ
- $k_B = 1$

Proton-proton ciklus



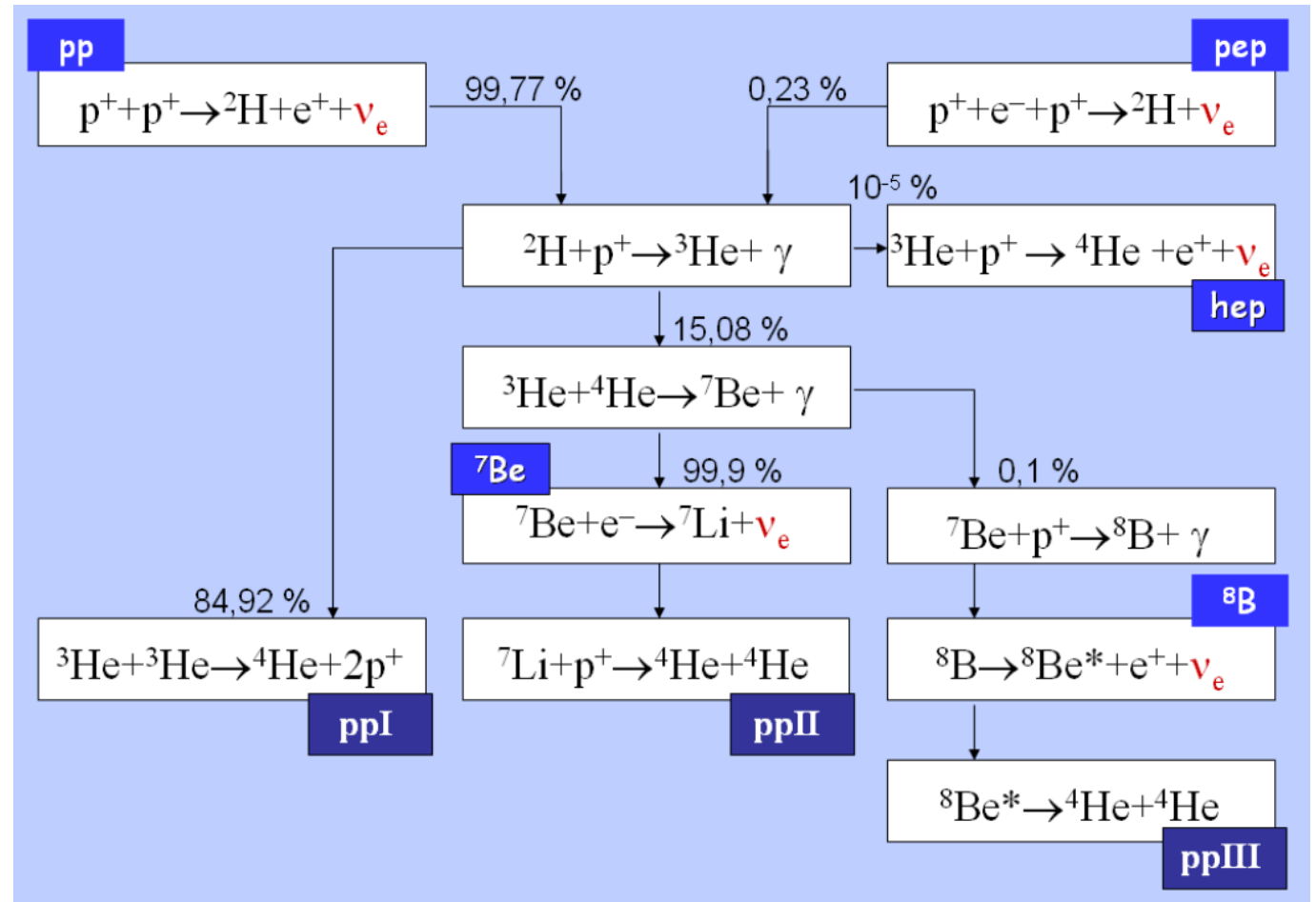
$$m_p = 938.272 \text{ MeV}/c^2$$

$$m_D = 1875.612 \text{ MeV}/c^2$$

$$M_e = 0.511 \text{ MeV}/c^2$$

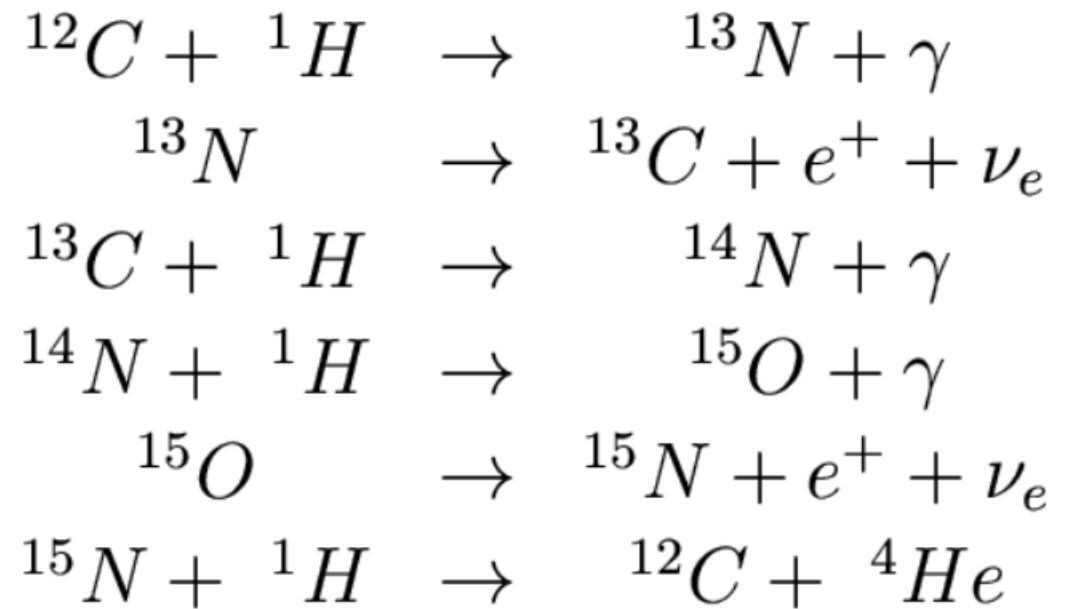
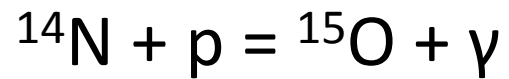
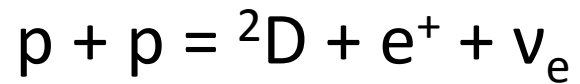
$$m_\nu < 0.12 \text{ eV}/c^2$$

$$c^2\Delta m = 0,421 \text{ MeV}$$



CNO-ciklus

Időskála:



Napállandó

A Nap által kisugárzott teljesítmény a Földpálya sugarában:

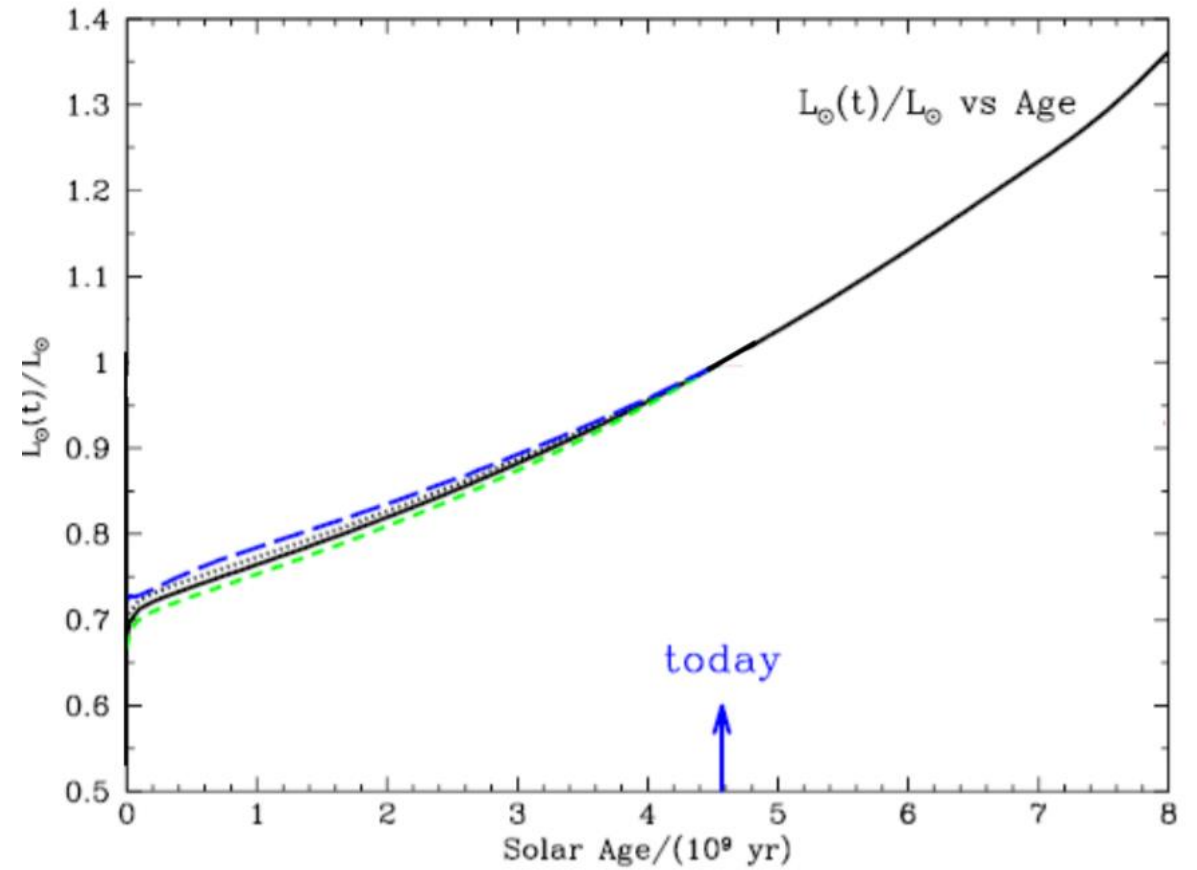
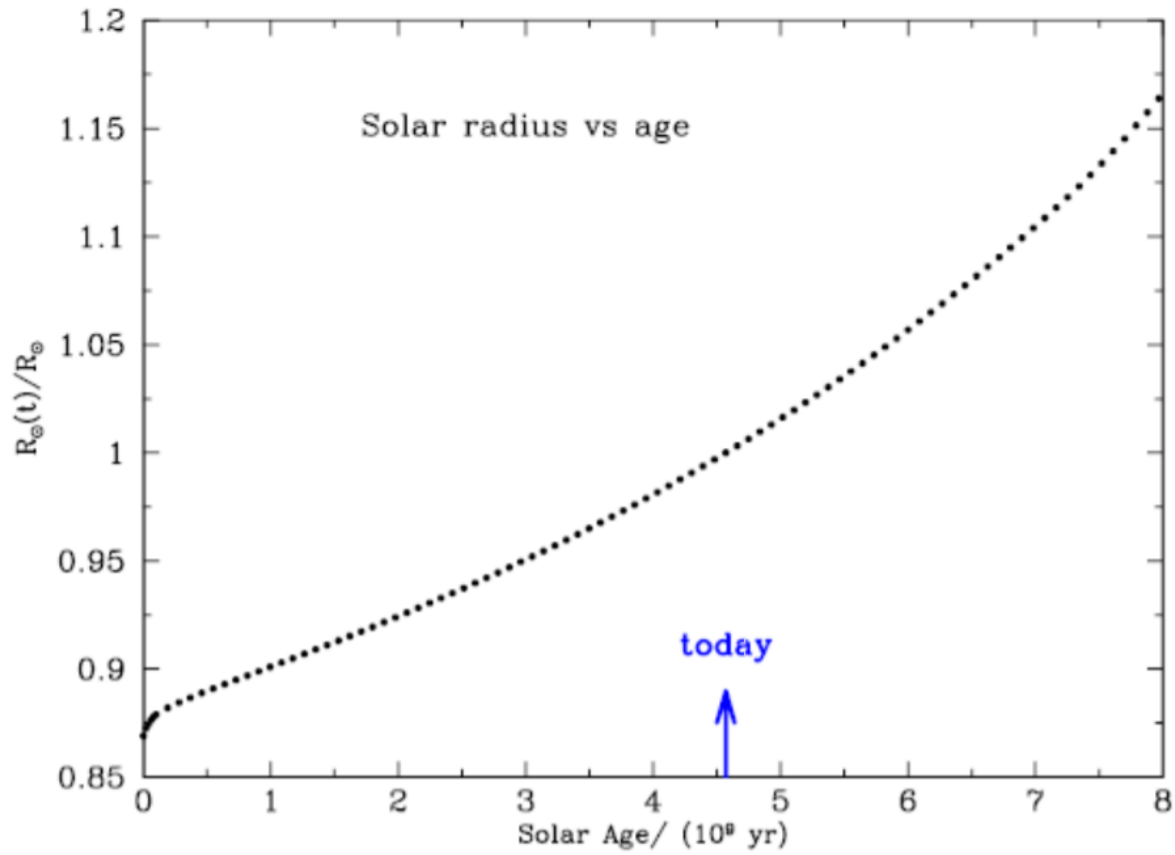
$$k = 1361 \text{ W/m}^2$$



$$k4\pi r^2 = \Delta m/\Delta t * c^2 \quad \longrightarrow \quad \Delta m/\Delta t = 7,71766 \text{ kg/s}$$

- Reakció ráta: $R = \frac{k4\pi r^2}{27MeV} = 10^{38} \frac{1}{s}$
- Neutrínó áram: $I = 2 \times 10^{38} \frac{\nu}{s}$
- Neutrínó fluxus: $\frac{I}{4\pi r^2} = 6 \times 10^{10} \frac{\nu}{cm^2 s}$

Feladatok:



Hivatkozások

- https://en.wikipedia.org/wiki/Solar_constant
- <http://matrix.elte.hu/~csoto/magfizszem/1.pdf>
- <http://matrix.elte.hu/~csoto/magfizszem/2.pdf>
- <http://matrix.elte.hu/~csoto/magfizszem/3.pdf>
- <http://matrix.elte.hu/~csoto/magfizszem/4.pdf>
- https://en.wikipedia.org/wiki/Stellar_structure
- <http://astro.u-szeged.hu/oktatas/asztrofizika/html/node25.html>
- K. Langanke & C. A. Barnes – Nucleosynthesis in the Big Bang and in Stars (World Scientific, Singapore, 1993)

Köszönöm a figyelmet!